

Cold-Electron Bolometers



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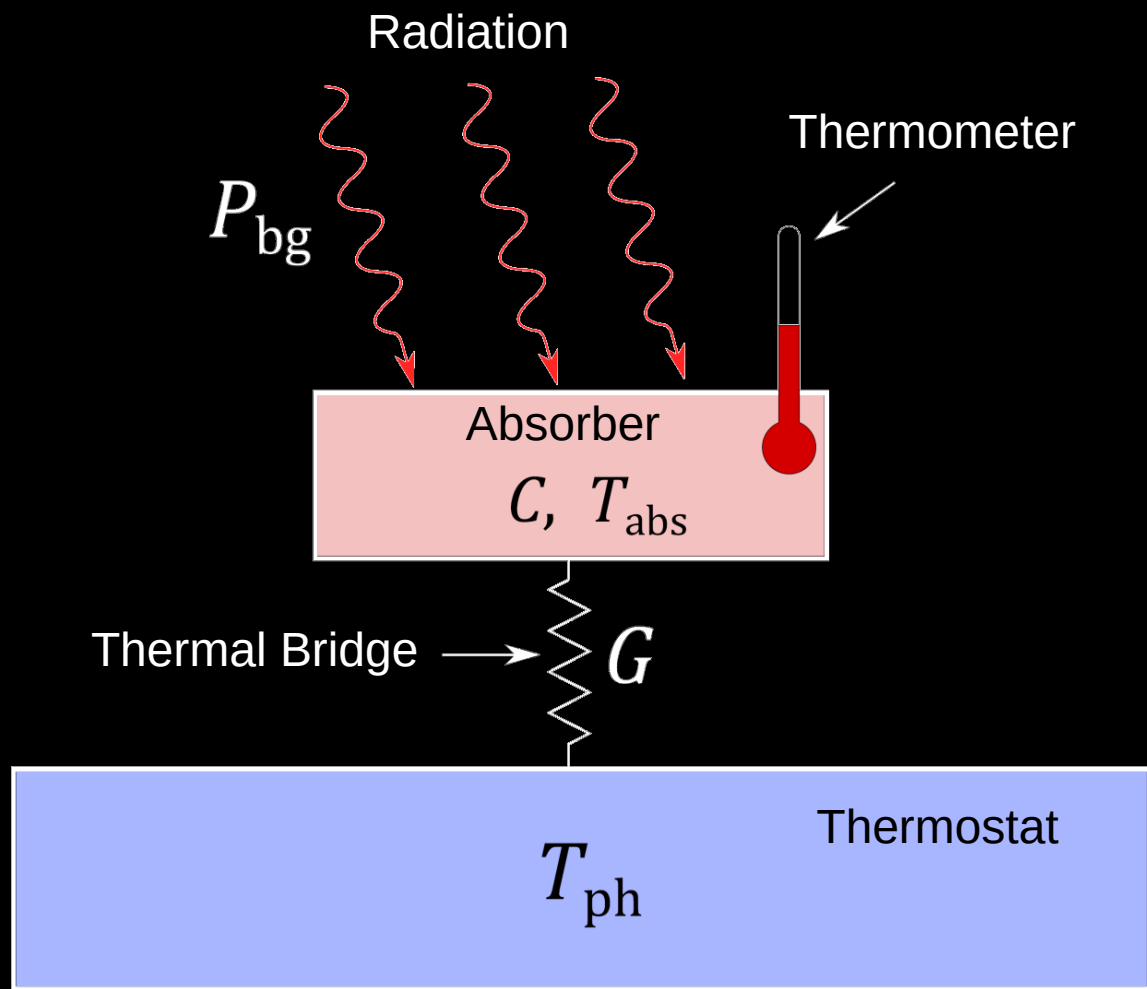
² Институт физики микроструктур РАН,

³ Chalmers tekniska högskola, Швеция,

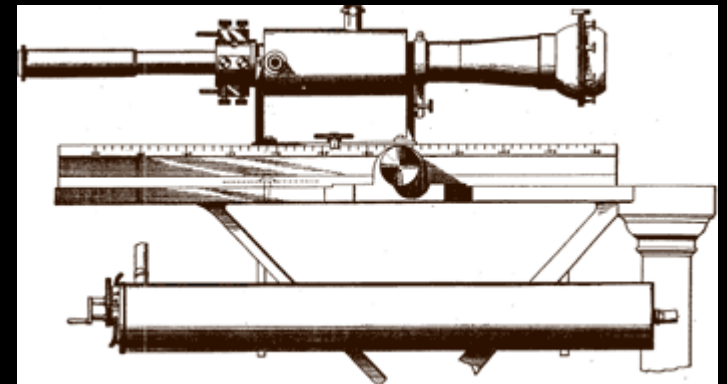
⁴ Московский институт электроники и математики им. А. Н. Тихонова,

⁵ Московский государственный университет им. М. В. Ломоносова.

Bolometers



$$\tau = C/G$$



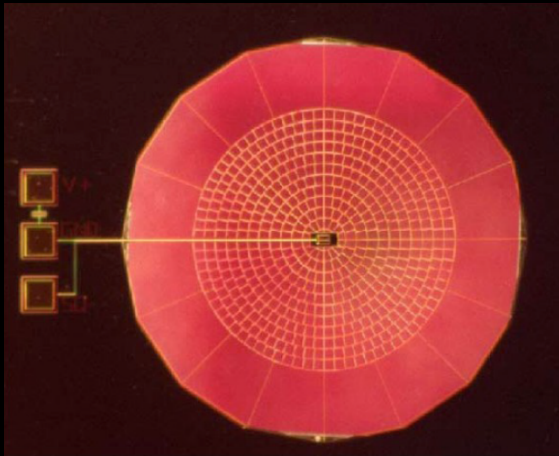
1878



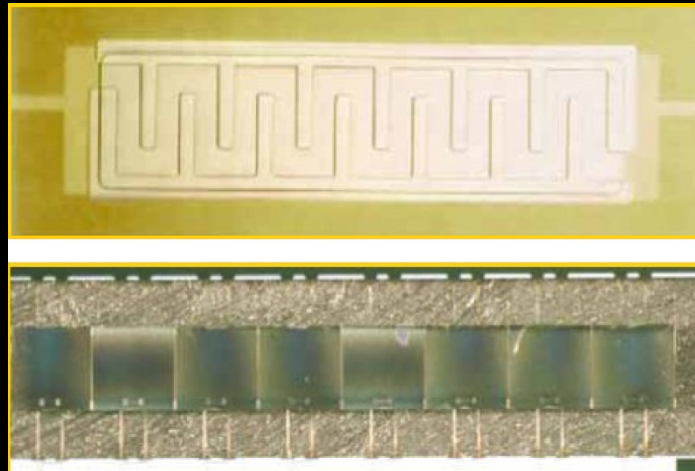
2016



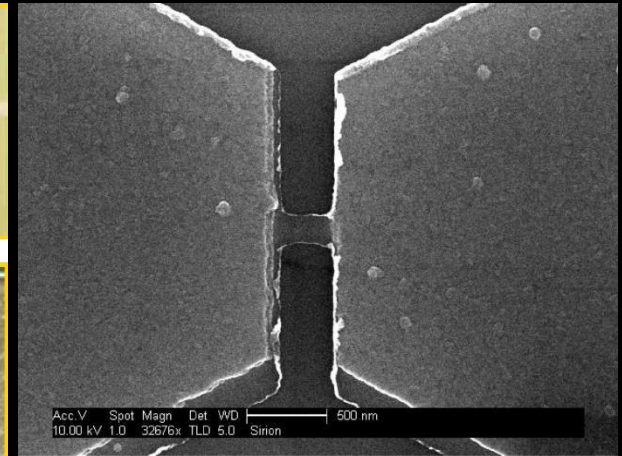
Bolometers



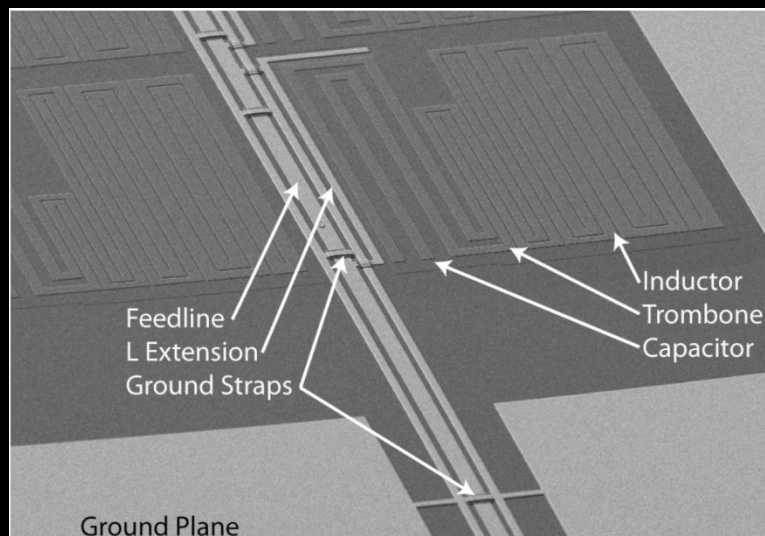
Semiconductor



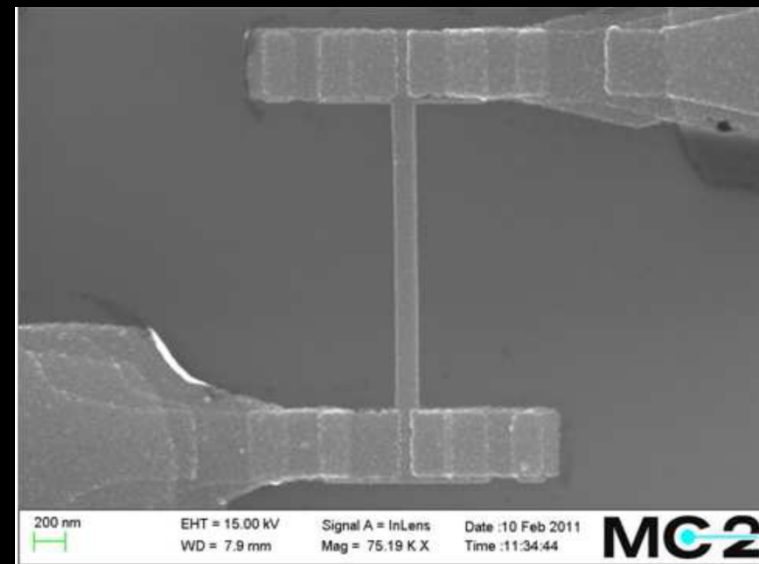
Transition Edge Sensor



Hot-Electron Bolometers

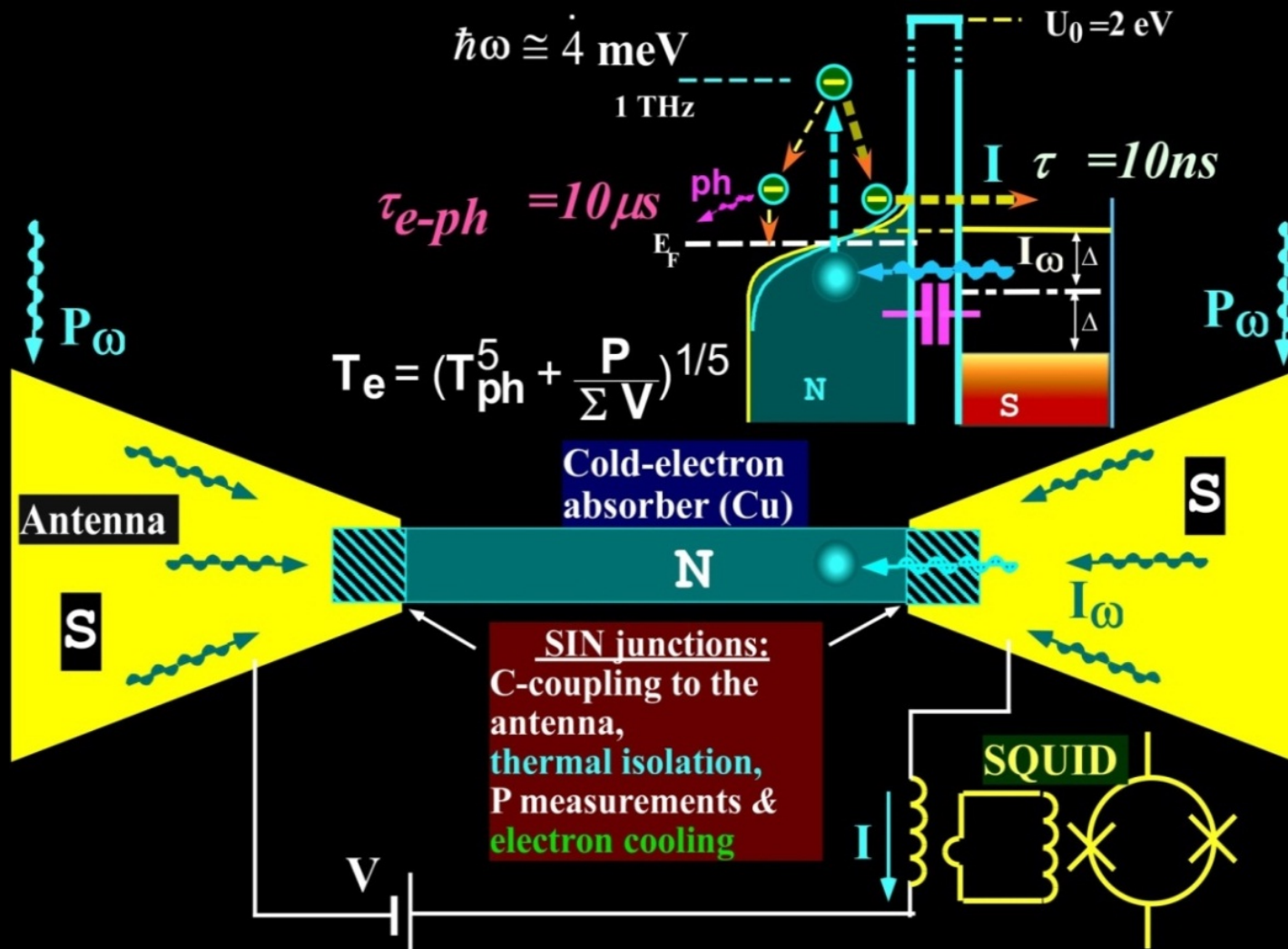


Kinetic Inductance Detector

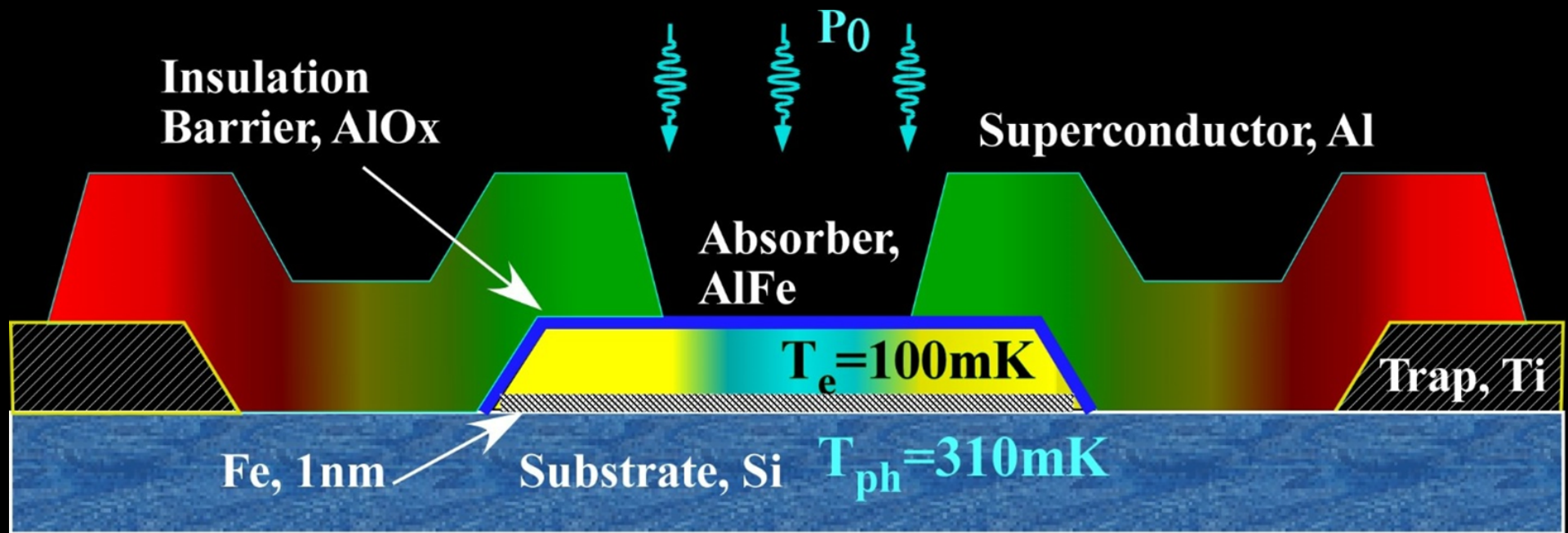


Cold-Electron Bolometers

Structure



Structure



Record Electron Cooling

Very small volume of the absorber (0.02 um^3) and large relative area of tunnel junctions (80% of the absorber).

Efficient removing of hot quasiparticles by double stocking them at both sides of the junctions.

Suppression of Andreev Reflection by a thin layer of Fe in AlFe absorber.

Result: unique bolometers operating at T_e less than T_{ph} .

Direct Electron Cooling vs Platform Phonon Cooling

Direct electron cooling

Kuzmin et al., 2002–2019

Direct electron self-cooling of the absorber at 300 mK to reach 70 mK or 170 mK under 20 pW radiation.

The cooler size is $1 \times 1 \text{ } \mu\text{m}^2$.

Result: photon noise limited system of 194 CEBs for OLIMPO Balloon Telescope.

Cooling Platform

Pekola et al. 1997–2019

Phonon cooling of a membrane from 300 to 200 mK.

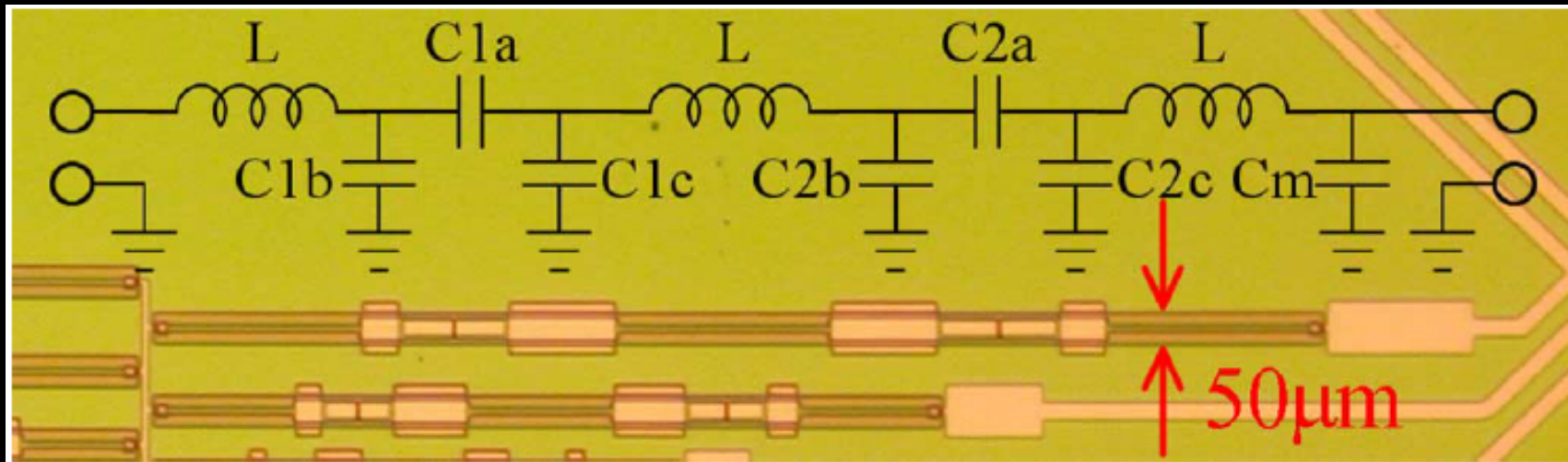
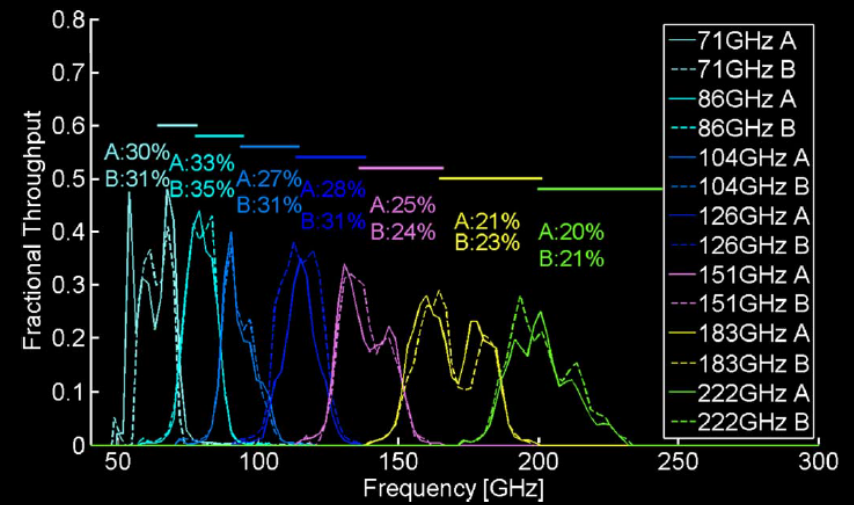
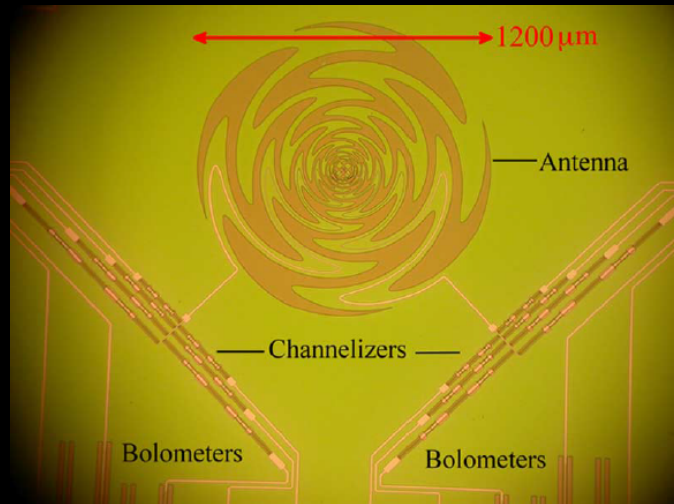
The cooler size is $2 \times 2 \text{ mm}^2$.

Result: no tests of bolometers on the platform so far.

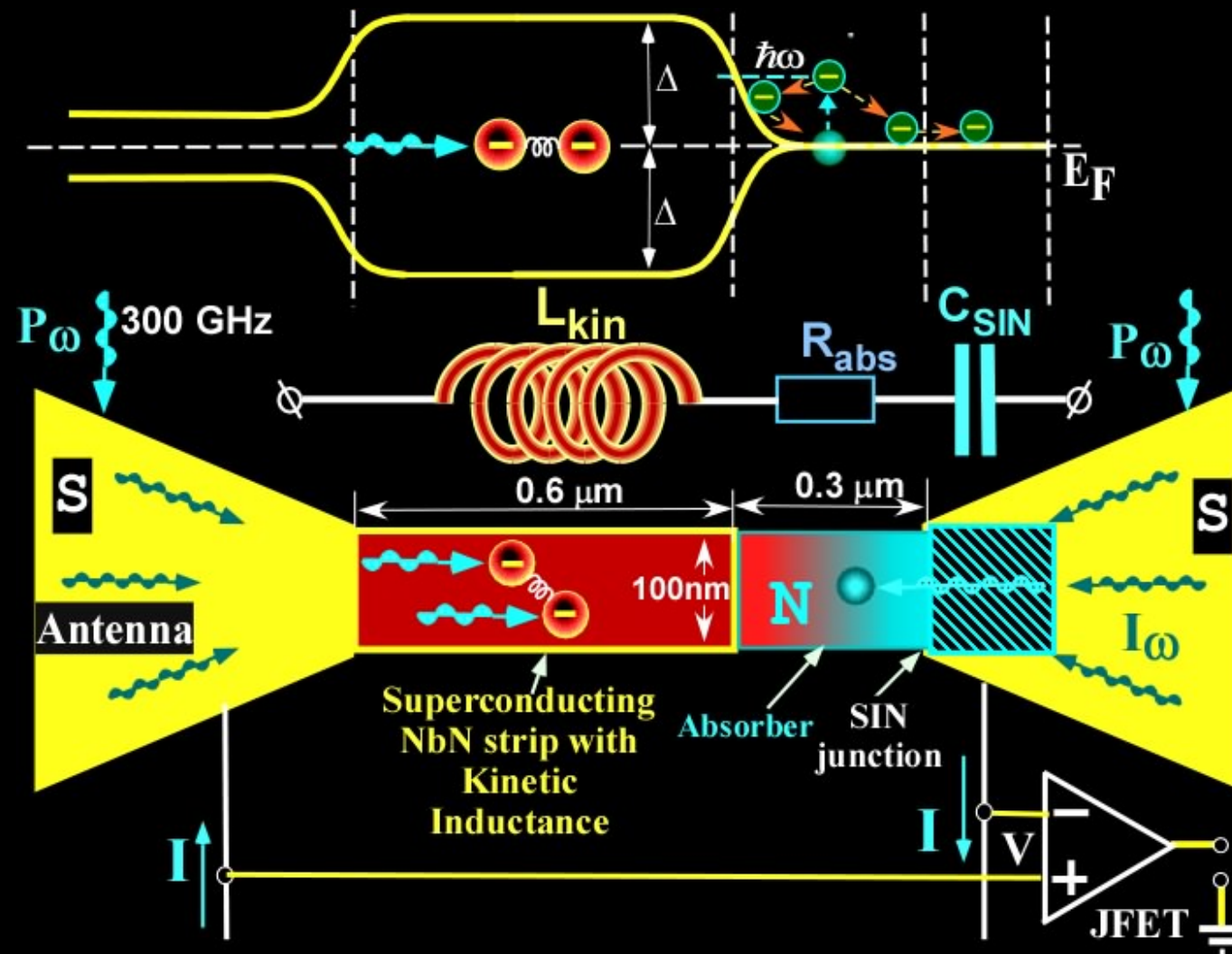
A. J. Manninen, M. M. Leivo, J. P. Pekola, “Refrigeration of a dielectric membrane by SIN tunneling”, Appl. Phys. Lett. 70, 1885 (1997).

Sinuuous Antenna

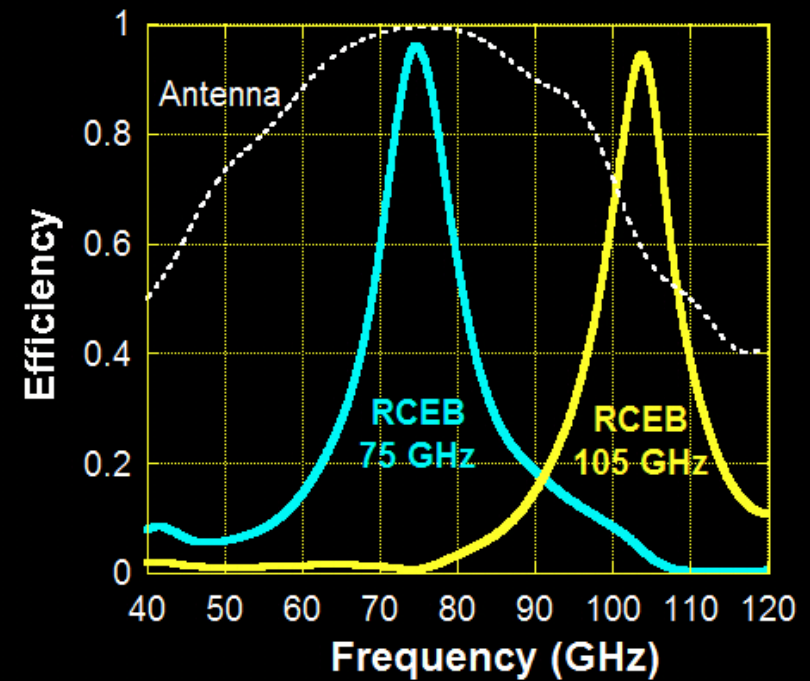
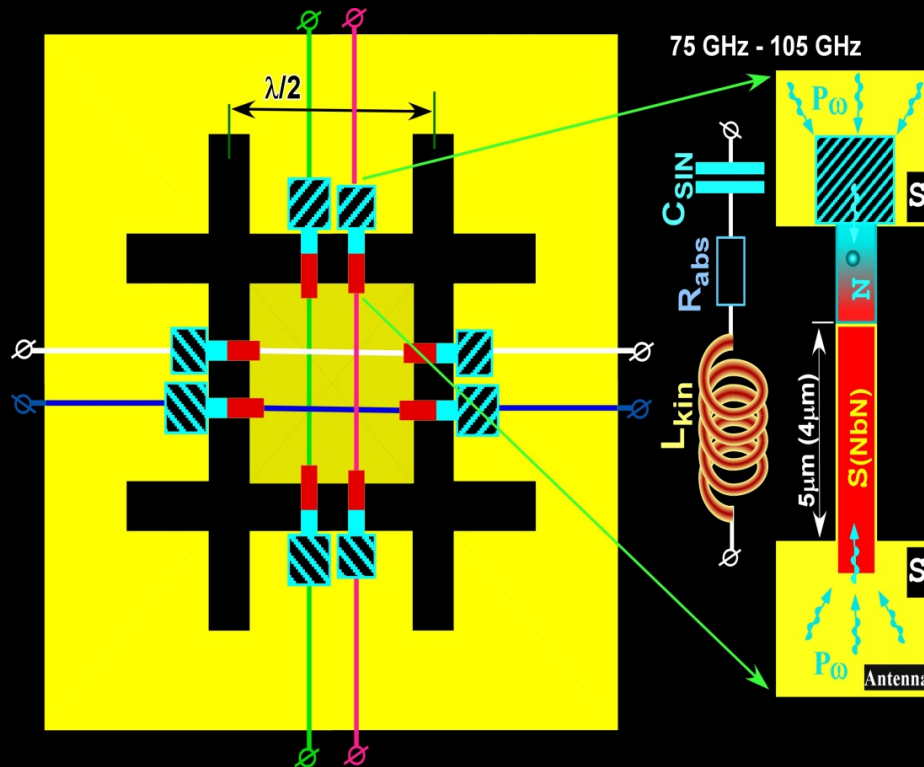
R. O'Brian *et al.*, IEEE Appl. Sc. (2011)



Nanofilter by a Kinetic Inductance of the NbN Strip

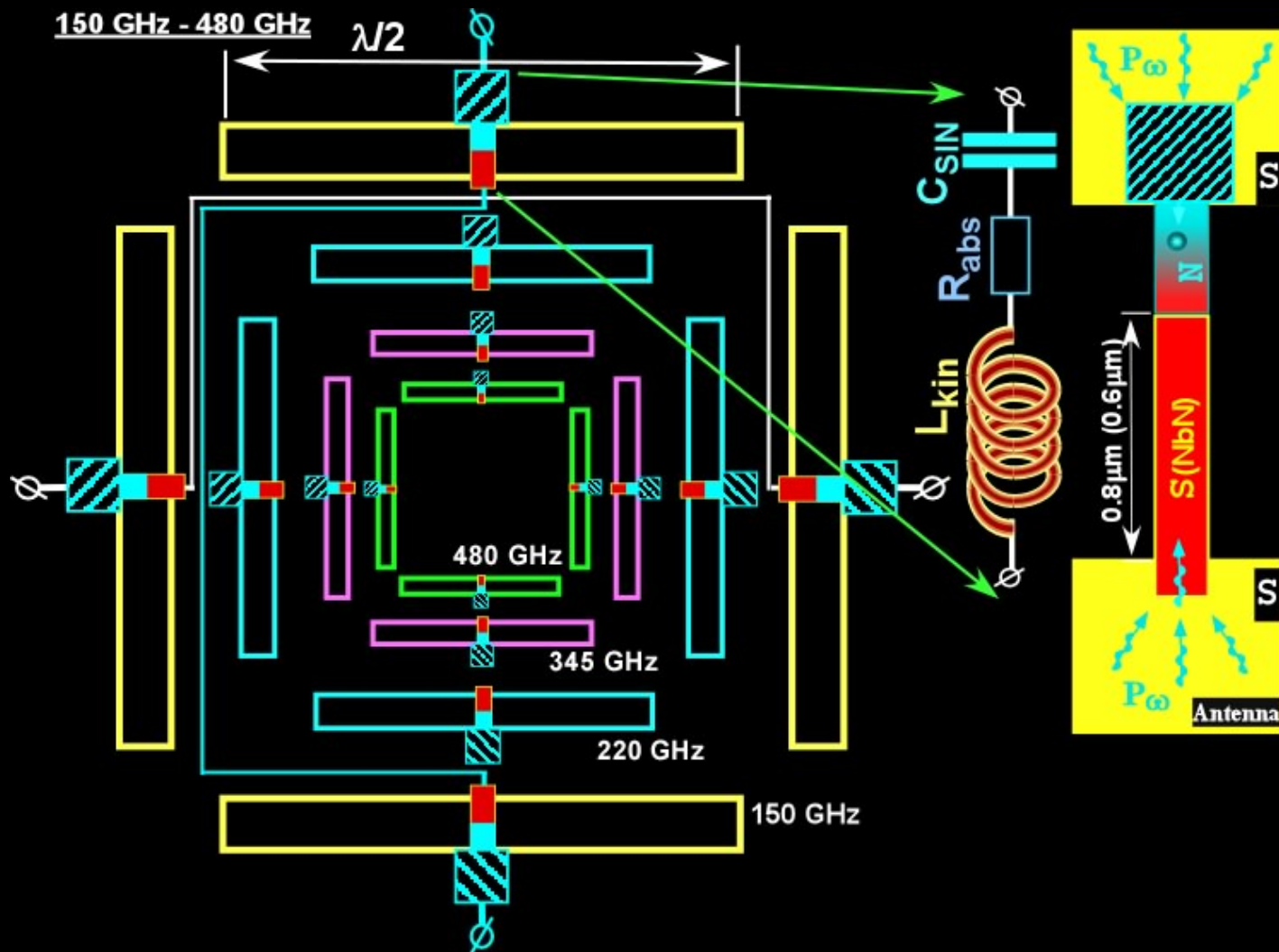


Cross-Slot Antenna

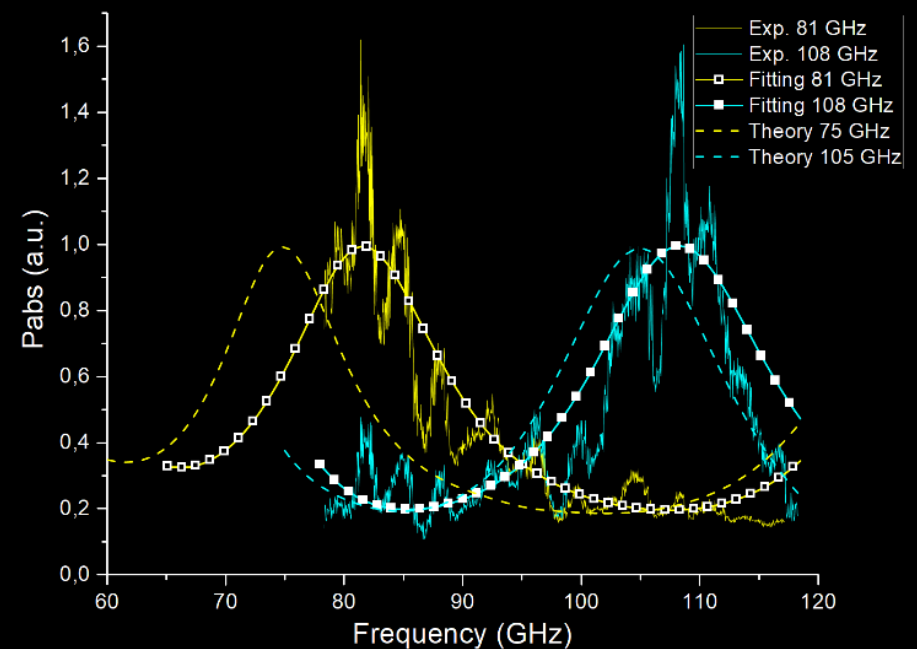
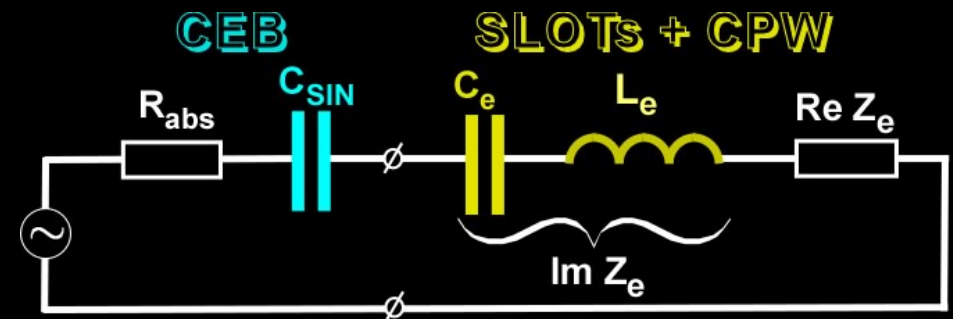


“Seashell” Slot Antenna

Leonid Kuzmin, Rome, 21 Sept 2013
stimulative discussions with Paolo de Bernardis are acknowledged



Seashell Antenna with Resonant Slot Antenna and w/o RCEBs



Application

MBR temperature fluctuations $\approx 160 \mu\text{K}$

Planck (2009–2013):

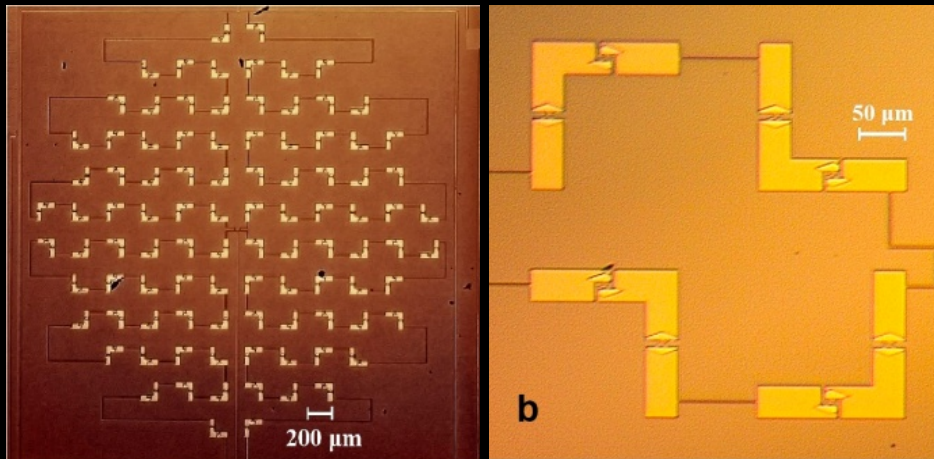
NEP $\approx 2 \times 10^{-17} \text{ W/Hz}^{1/2}$, 100 mK

Estimated upper B-mode limit $\approx 0.1 \mu\text{K}$

COrE (2015–2025):

NEP $\approx 10^{-18} \text{ W/Hz}^{1/2}$, 100 mK

Application: OLIMPO Receiver Pixel

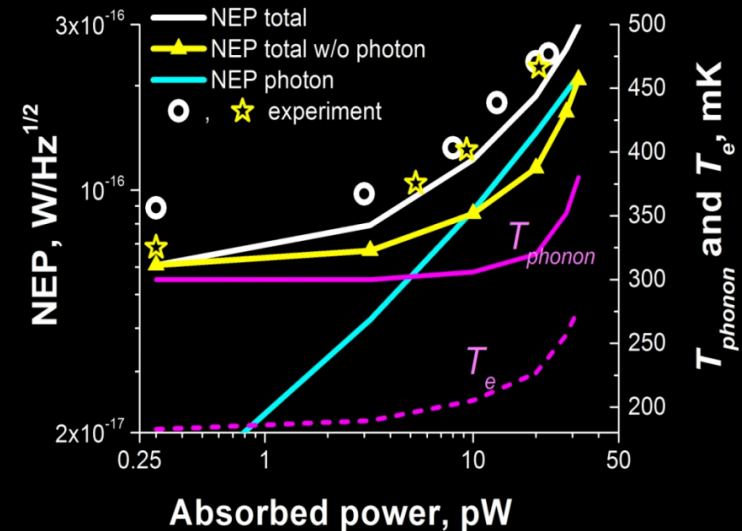
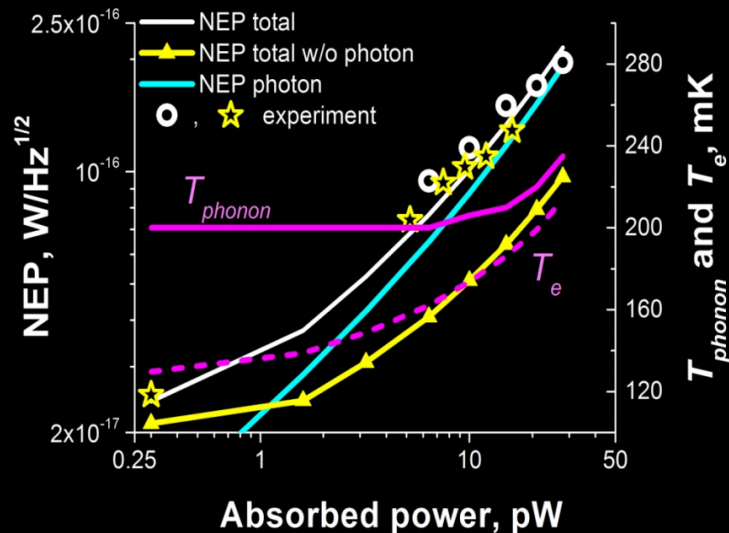


OLIMPO, channel III:
350 GHz, $\Delta f = 33$ GHz

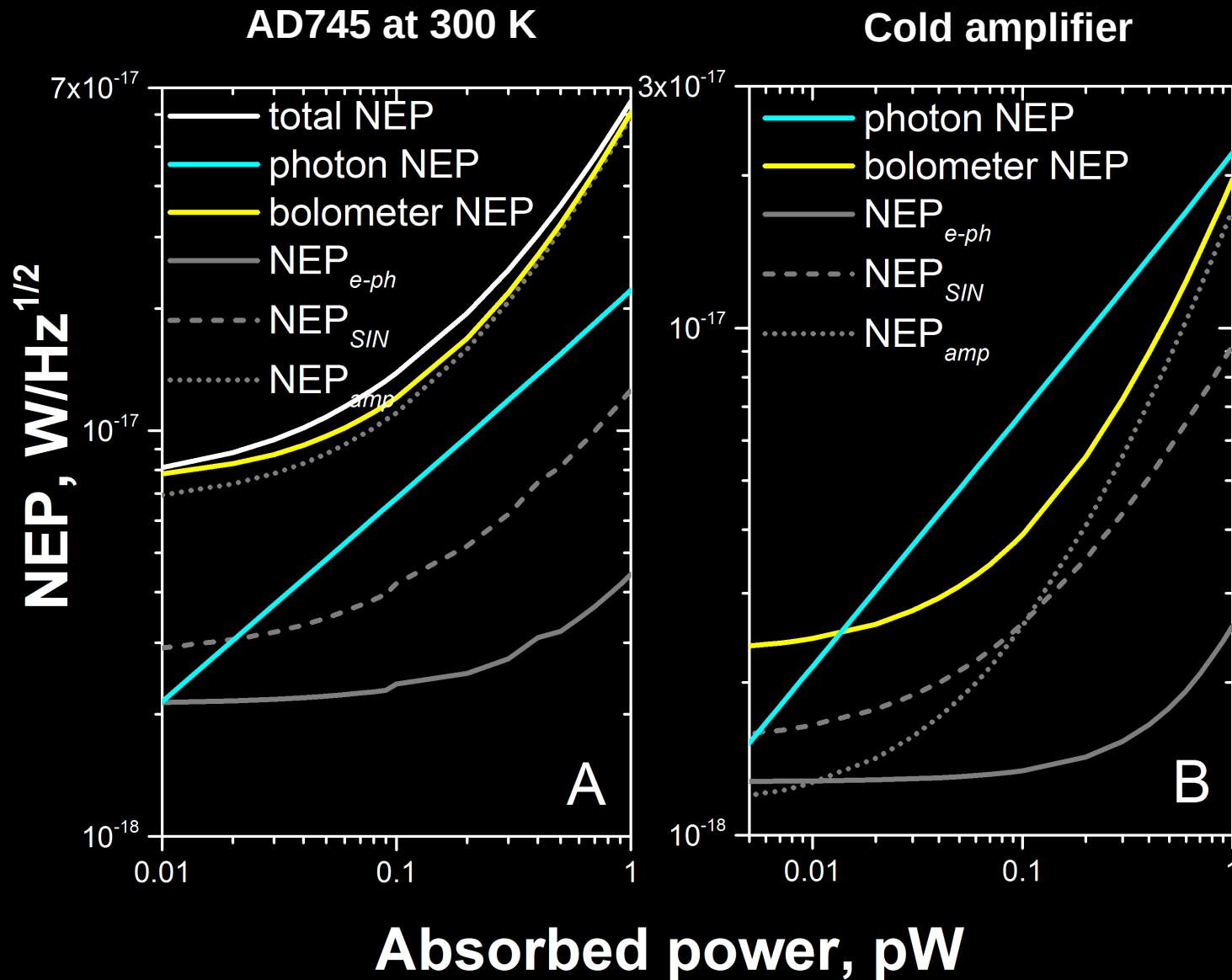
Load: $P = 38$ to 66 pW

Photon noise limit:

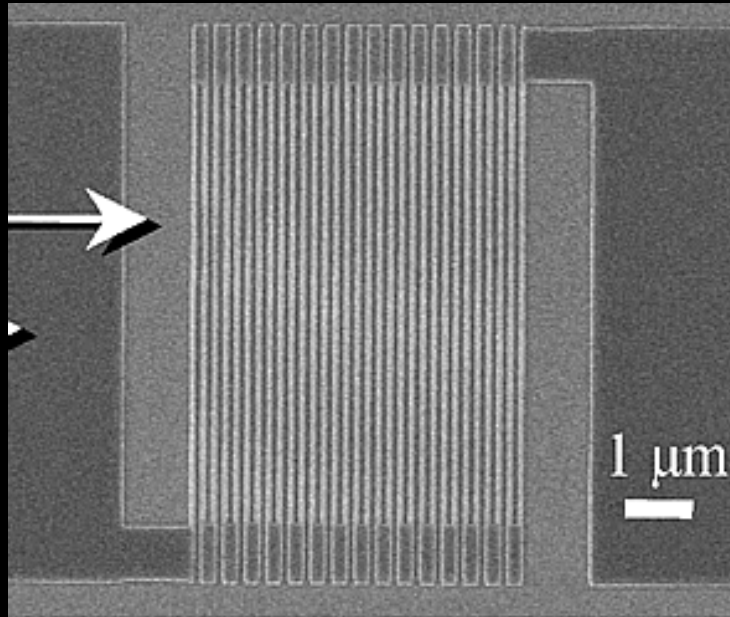
$$NEP_{bol} < NEP_{ph} = 10^{-16} \text{ W/Hz}^{1/2}$$



Single Bolometer NEP

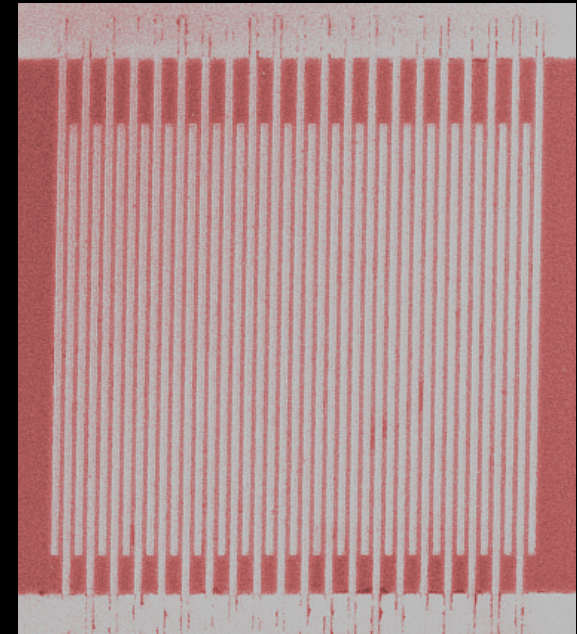


Weak Signals Detector



SEM of a nanowire pattern

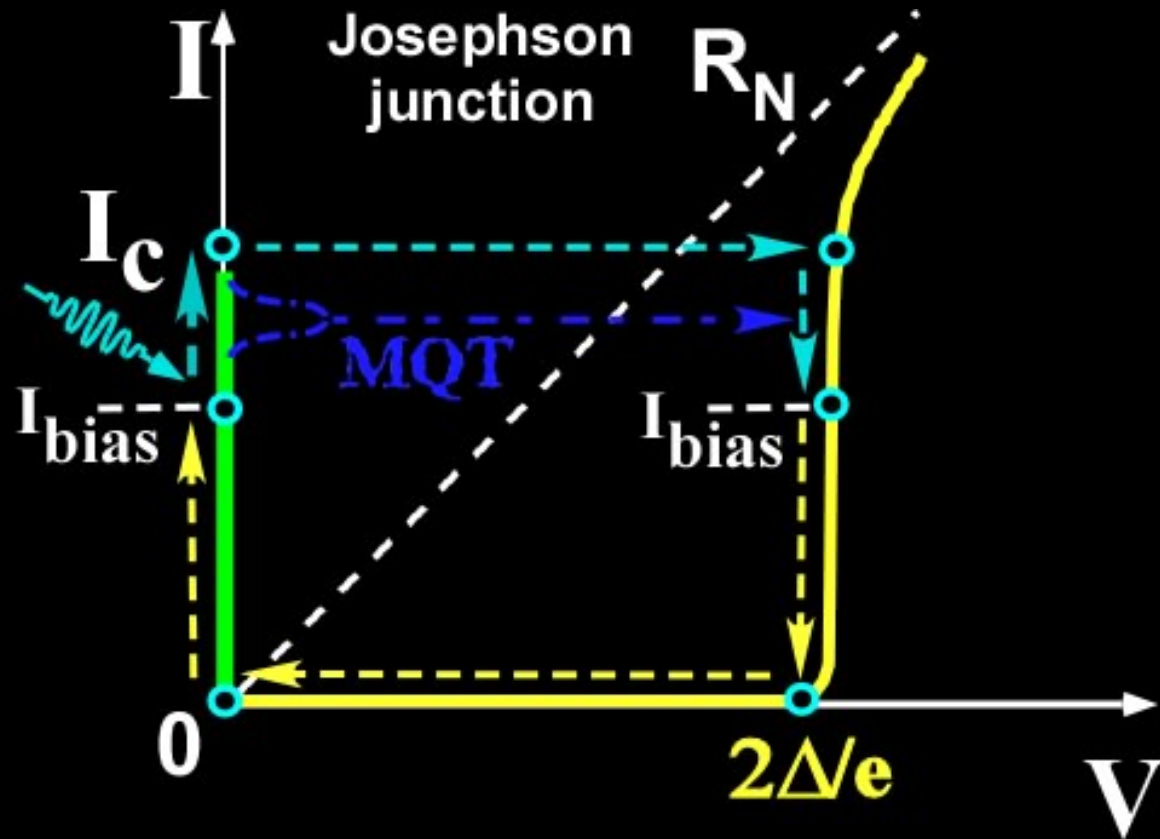
G. Gol'tsman et al.



WSi nanowire with a detection energy threshold of 0.8 eV

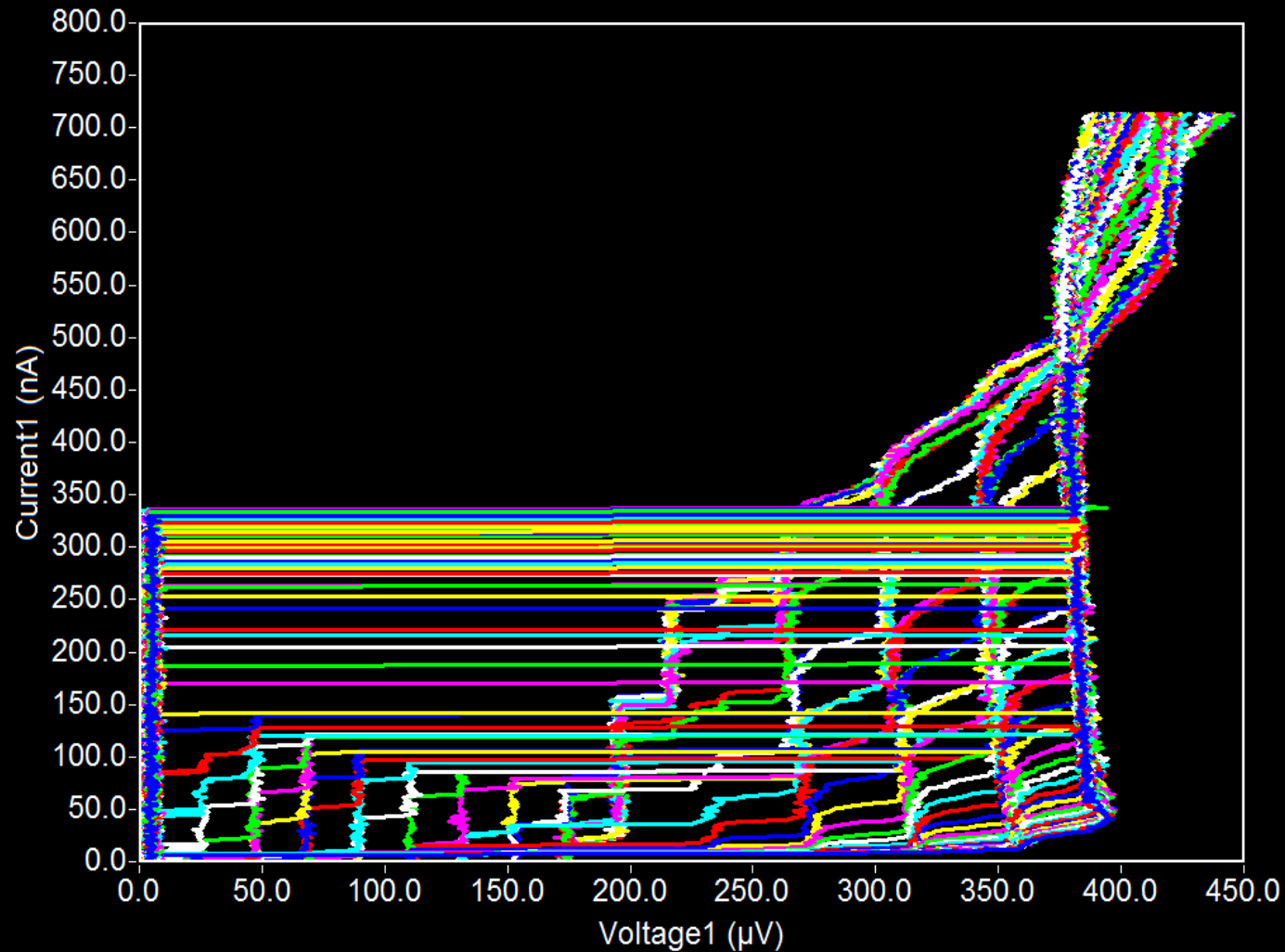
Y. Hochberg et al.

Weak Signals Detector



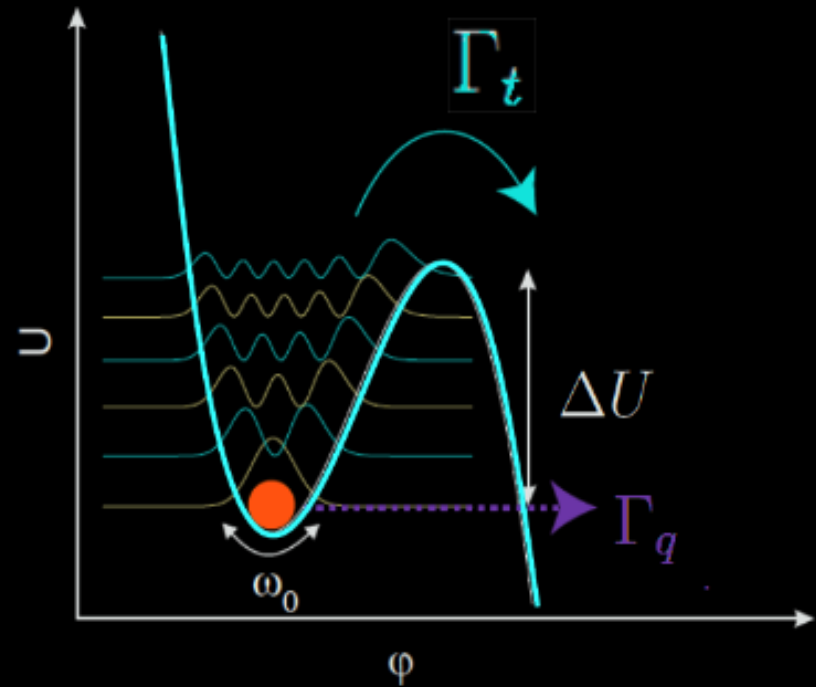
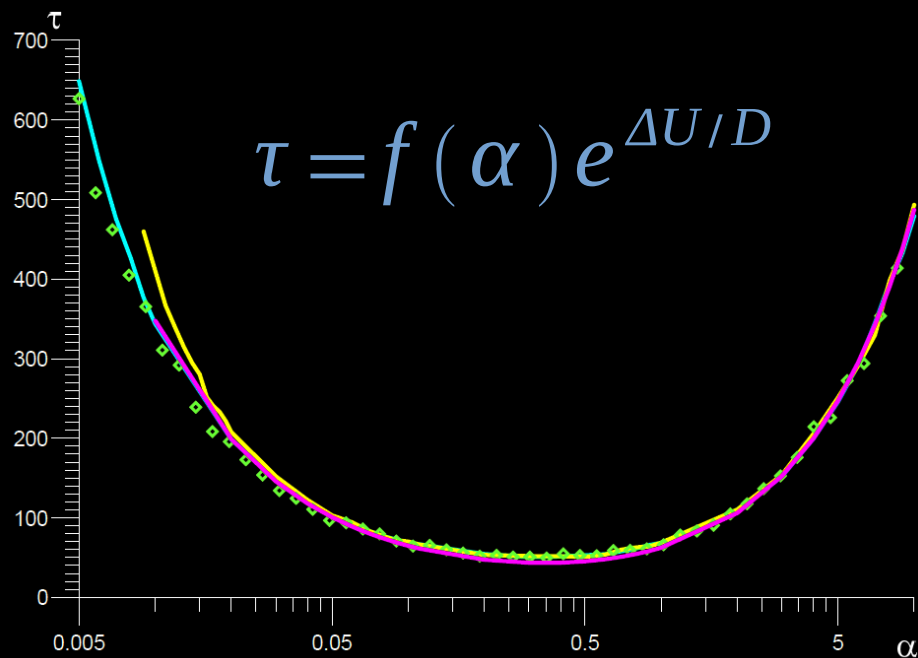
G. Oelsner, L. S. Revin, E. Il'ichev, A. L. Pankratov, H.-G. Meyer, L. Gronberg, J. Hassel, L. S. Kuzmin, Appl. Phys. Lett. 103, 142605 (2013)

9 GHz Signal Response

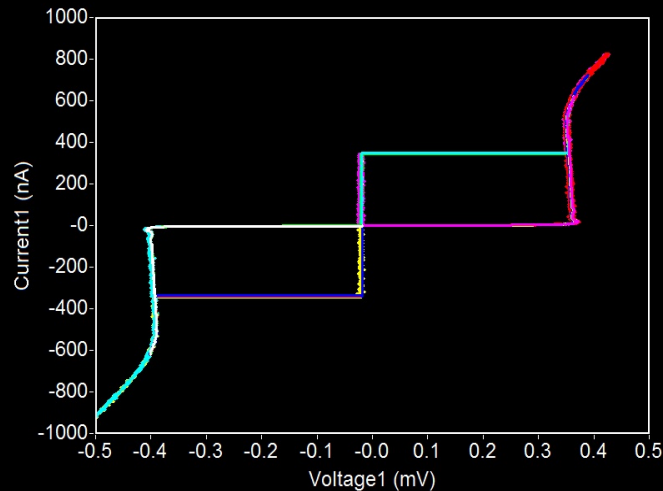


Theory

$$\frac{d^2 \varphi}{dt^2} + \alpha \frac{d \varphi}{dt} = \eta - \sin(\varphi) + \sqrt{2D} \xi(t)$$



Измерение ВАХ образцов Al SIS контактов



SIS1

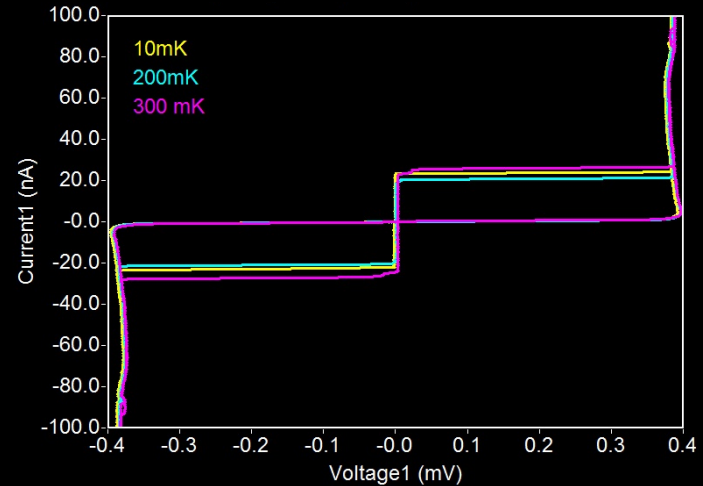
$$I_c = 0.328 \text{ } \mu\text{A}$$

$$S = 7 \text{ } \mu\text{m}^2$$

$$C = 0.315 \text{ pF}$$

$$R_n = 470 \text{ Ohm}$$

$$\alpha = 0.1$$



SIS2

$$I_c = 0.025 \text{ } \mu\text{A}$$

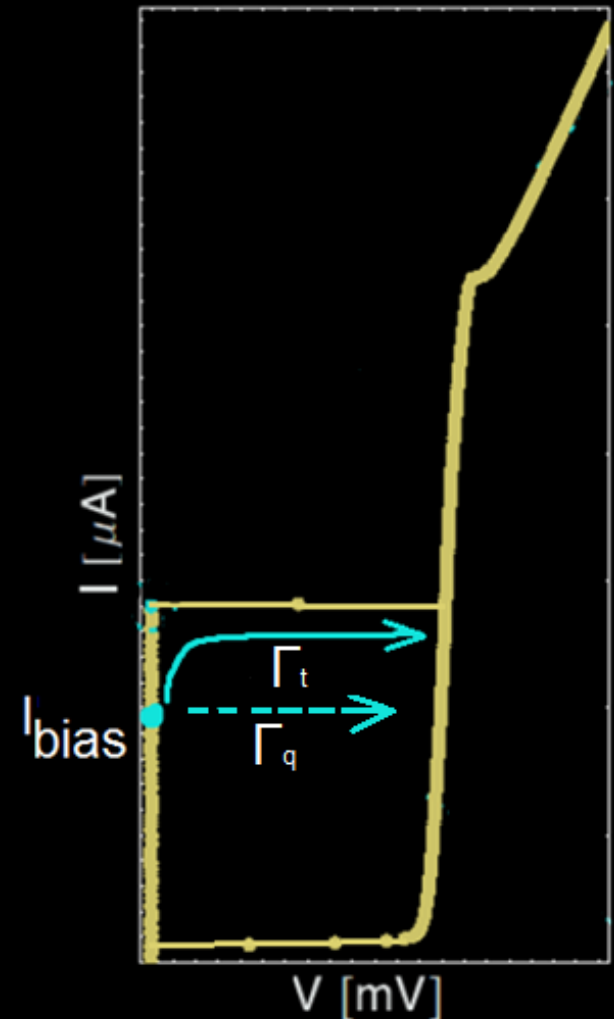
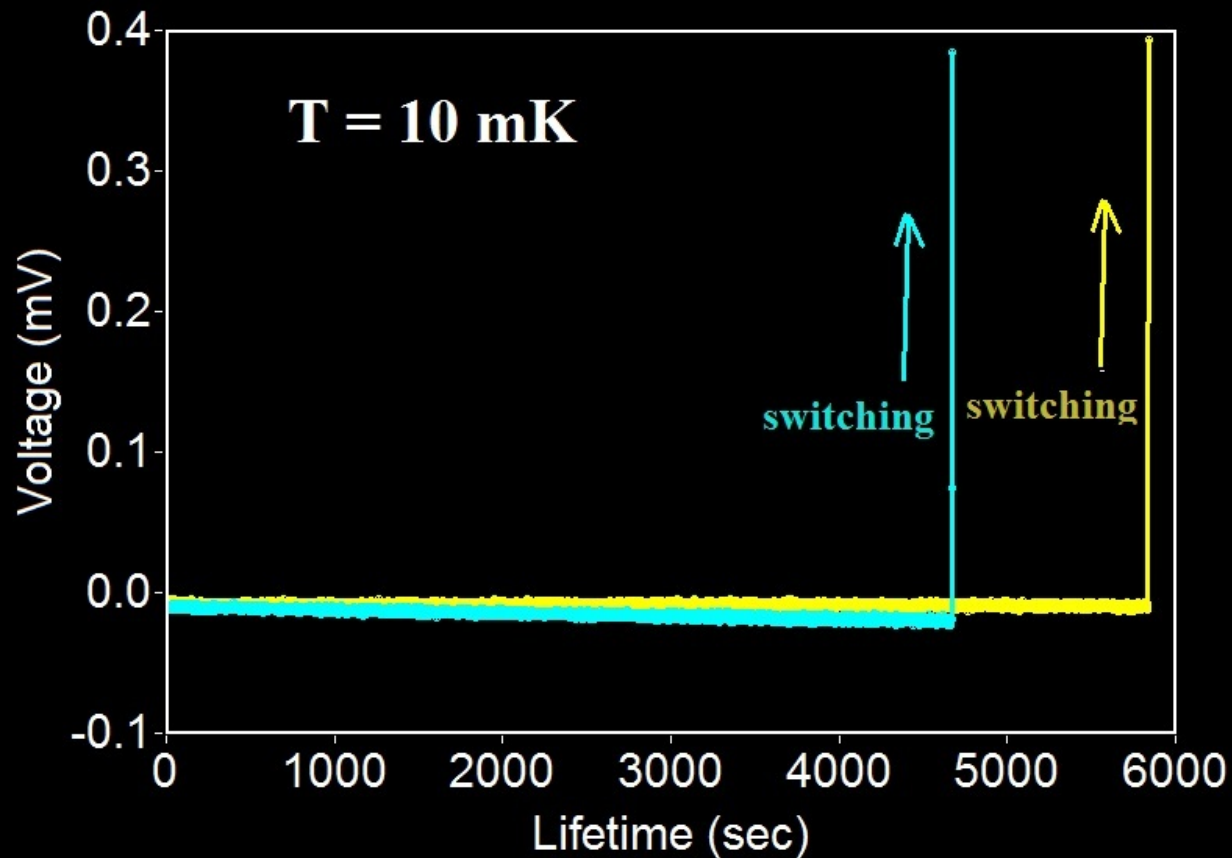
$$S = 0.8 \text{ } \mu\text{m}^2$$

$$C = 0.036 \text{ pF}$$

$$R_n = 2000 \text{ Ohm}$$

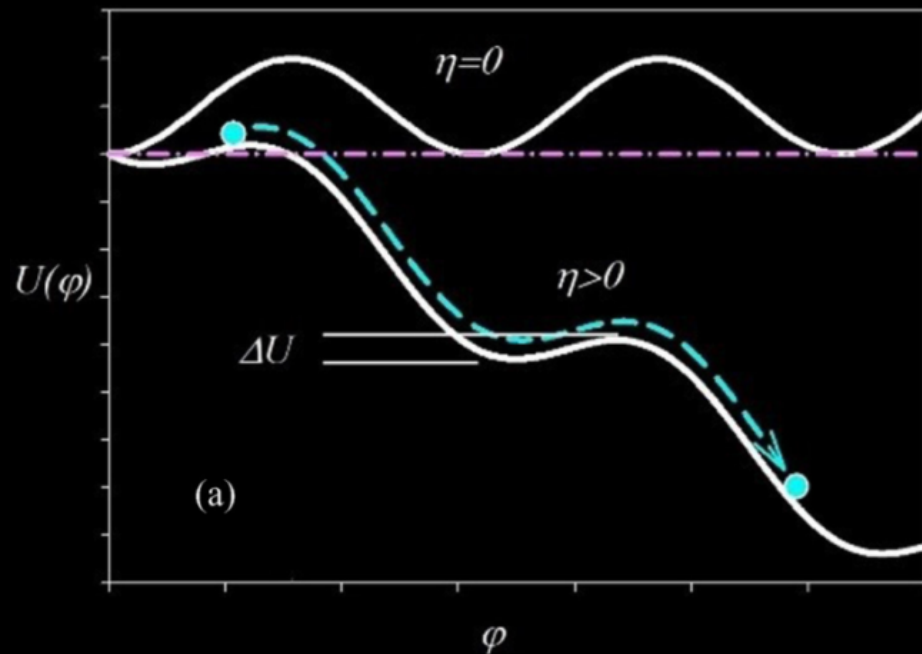
$$\alpha = 0.3$$

Lifetime: Dark Counts Rate

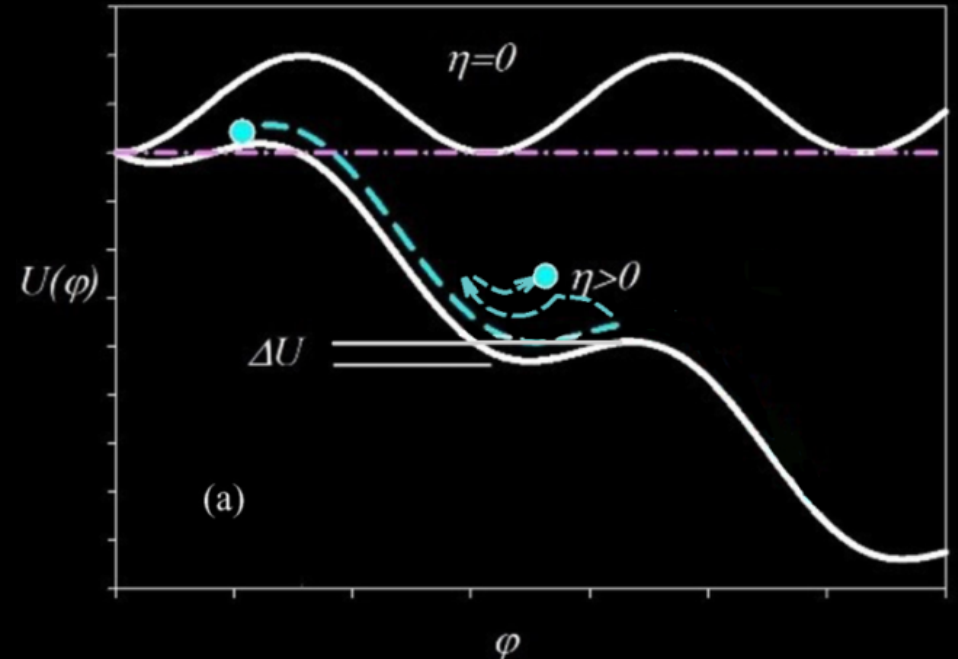


Running State vs Phase Diffusion State

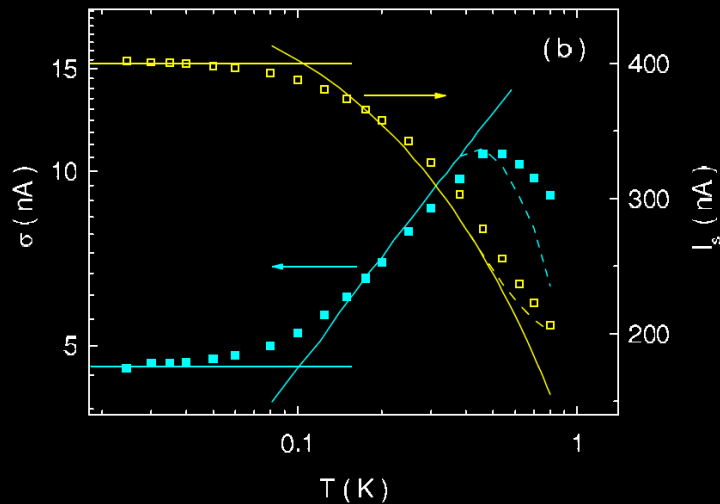
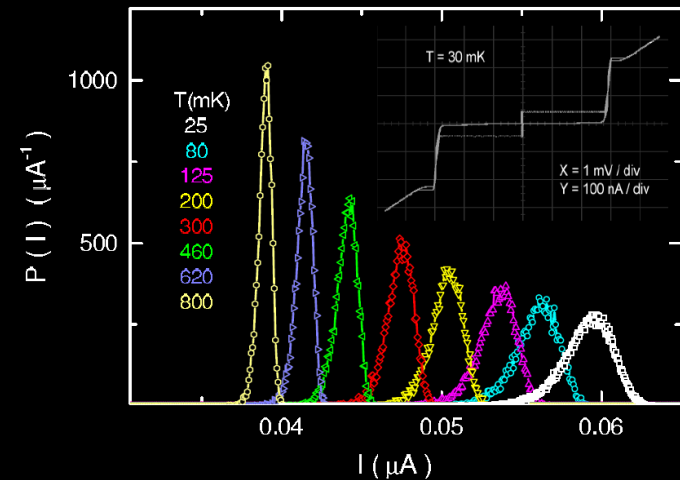
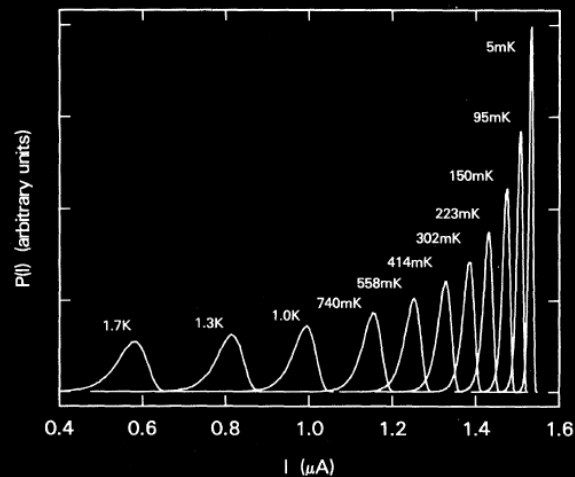
Running State:



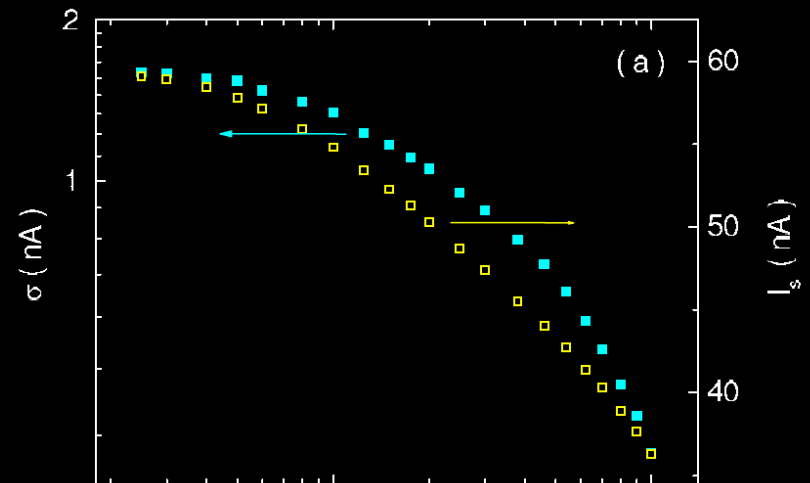
Phase Diffusion:



Running State vs Phase Diffusion State

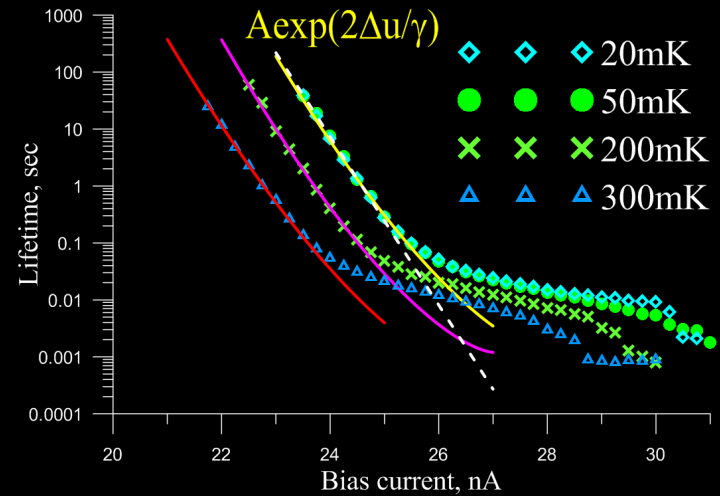
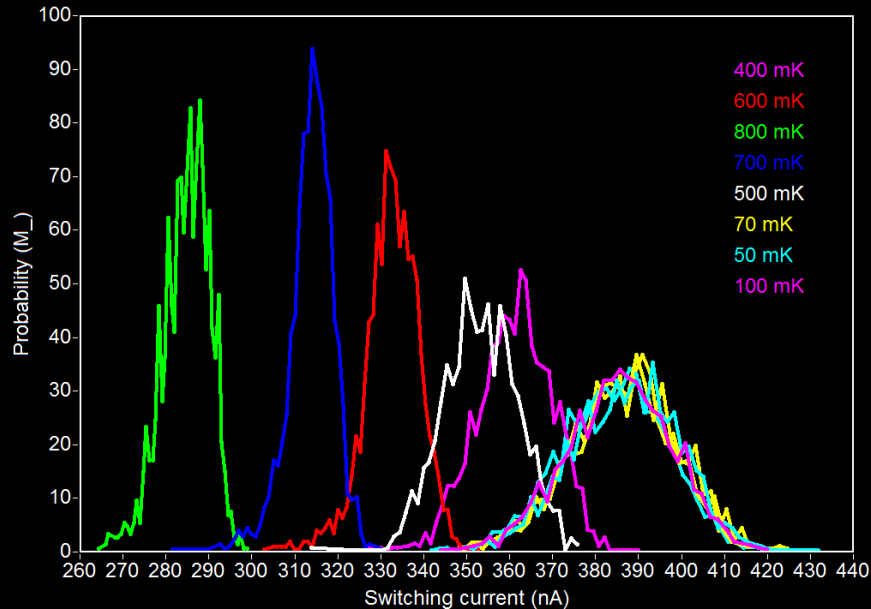


Running State



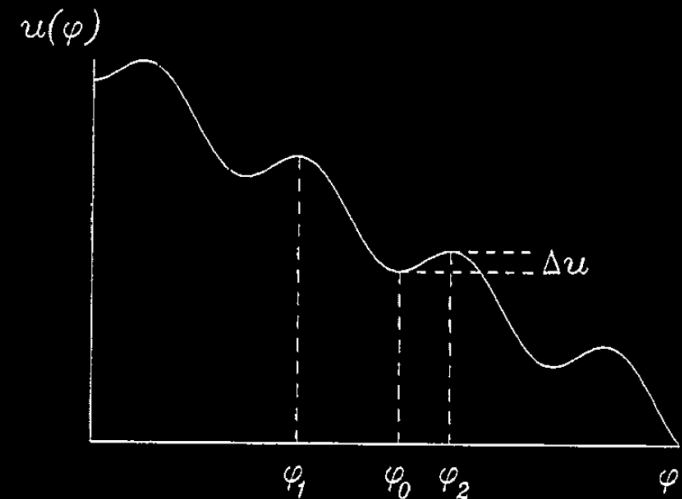
Phase Diffusion

Lifetime Distribution

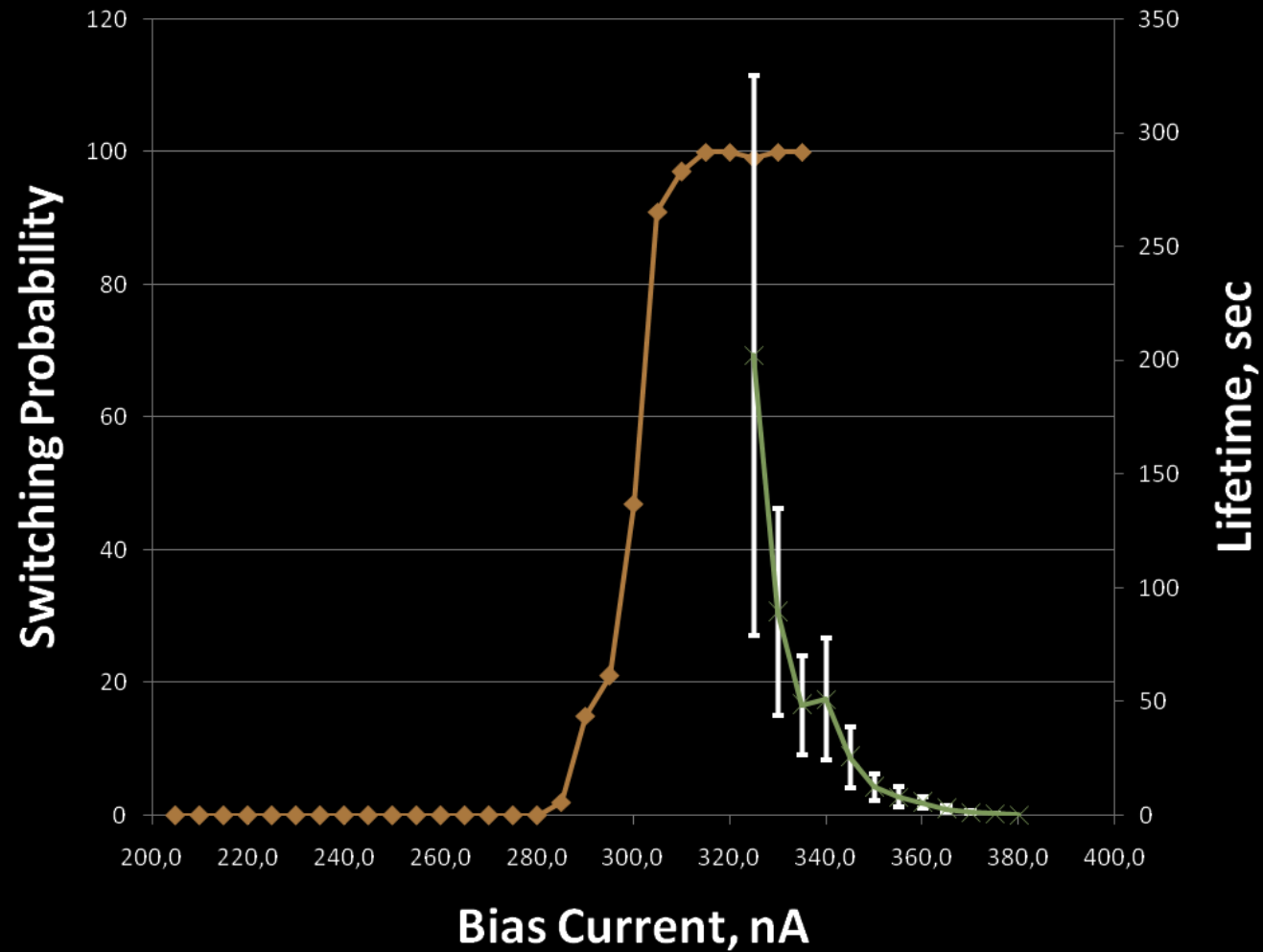


$$\Delta u / \gamma = (2\sqrt{1 - i^2} + 2i(\arcsin i - \pi/2)) / \gamma$$

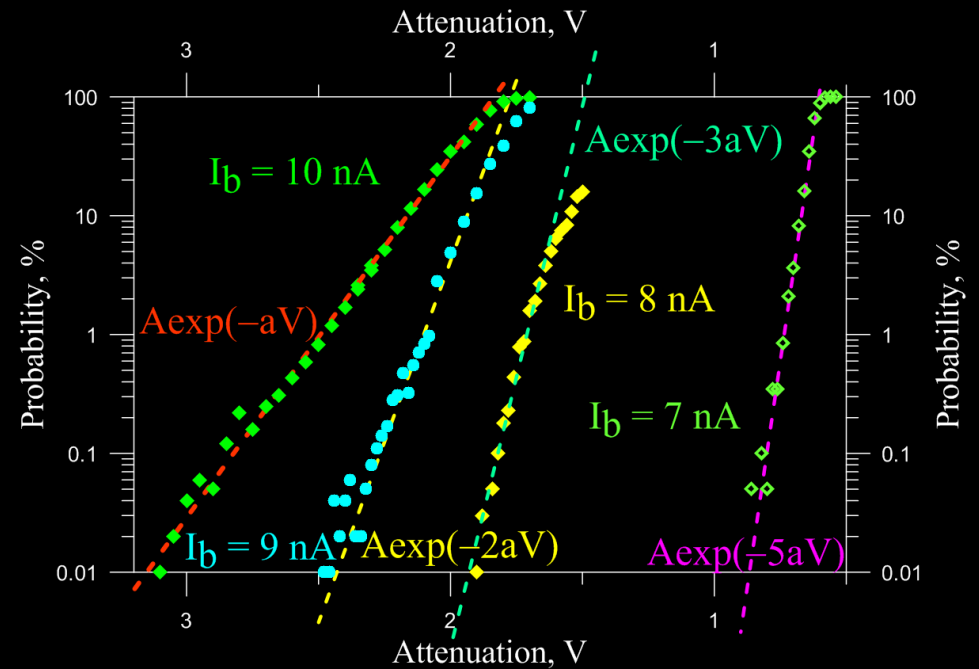
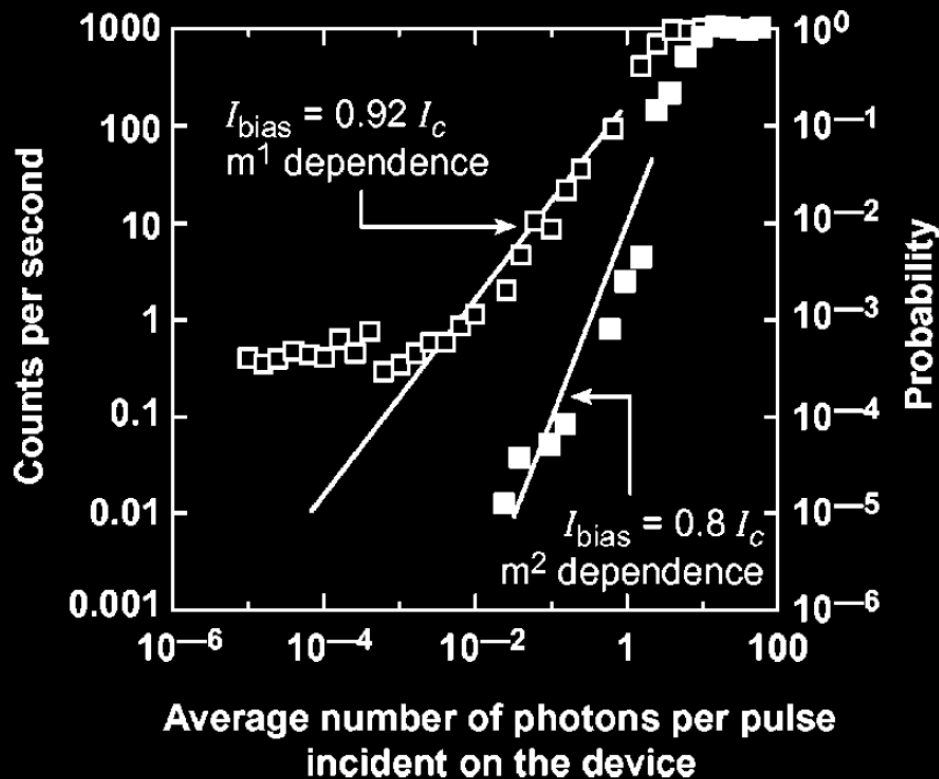
$$\gamma = 2ekT / \hbar I_c = I_T / I_c$$



9 GHz Signal Response



9 GHz Signal Response



G. N. Gol'tsman, O. Okunev, G. Chulkova, A. Lipatov, A. Semenov, K. Smirnov, B. Voronov, and A. Dzardanov, Appl. Phys. Lett., 79, 705 (2001).

Conclusions

- A detector of an ultimate sensitivity has been created.
- There are many more mysteries to deal with.